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Objective

Develop a massively scalable Monte Carlo radiation (SPARTA-rMC) code to simulate radiative transport through real complex materials.

Introduction

Space vehicles entering a planet's atmosphere attain hypersonic speeds (M>10). At such high-speeds, extreme thermal conditions are experienced and a TPS material is required to protect the vehicle from overheating.



Figure: 1. Atmosphere Re-entry of a space vehicle

It had been previously believed that the incoming shock-layer radiative flux was either reflected at the surface or absorbed within a 1 mm thickness of the surface [4]. Since most heritage TPS are far thicker than this 1 mm absorption zone, the penetration of radiation within the TPS has always been neglected. However, new experimental work shows that spectral radiation can penetrate the TPS and affect the material response over a significant depth, with higher intensity emissions penetrating deeper into the TPS [5]. As the temperature across the TPS increases, the thermal decomposition (pyrolysis) of the phenolic resin creates a porous network and increased penetration is observed in the decomposed (char) zone. Understanding radiative transport is therefore, essential to designing a reliable TPS material.

A Numerical Approach to Calculate In-Depth Penetration of **Radiative Heating in Thermal Protection System (TPS)**

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Theory

In a Monte Carlo radiation code, the photons are A geometric optics approach has been implemented tracked using the ray tracing approach. When the in the Sparta-rMC code. Once the code imports the photon encounters an object in its path, a decision microstructure geometry then photons are generated tree determines scattering, transmission, or absorpwithin the domain. Using the Monte Carlo ray trace method, the extinction and absorption coefficients tion depending on the probabilities and efficiencies are calculated by summing the extinction length and determined from quantum and classical mechanics the absorption length computed from the photon's theory and the event occurrences are determined mean free path [1]. stochastically.

Central Idea

Developing a microscale approach to compute optical properties of multi-phase material microstructures, followed by a macroscale analysis of radiative transport.

Figure: 2. Microstructure of LI 900 TPS Material

 \rightarrow LI 900 is a TPS material which is essentially Silicon Dioxide. It consists of a fibrous network of infinitely long cylinders with diameter of 5 micron with variable standard deviation and porosity 0.94.

Results

 \rightarrow We have compared our work with the results of Marschall [2]. \rightarrow With confidence in our results, we move forward to generate the optical properties for different TPS material.

Methods

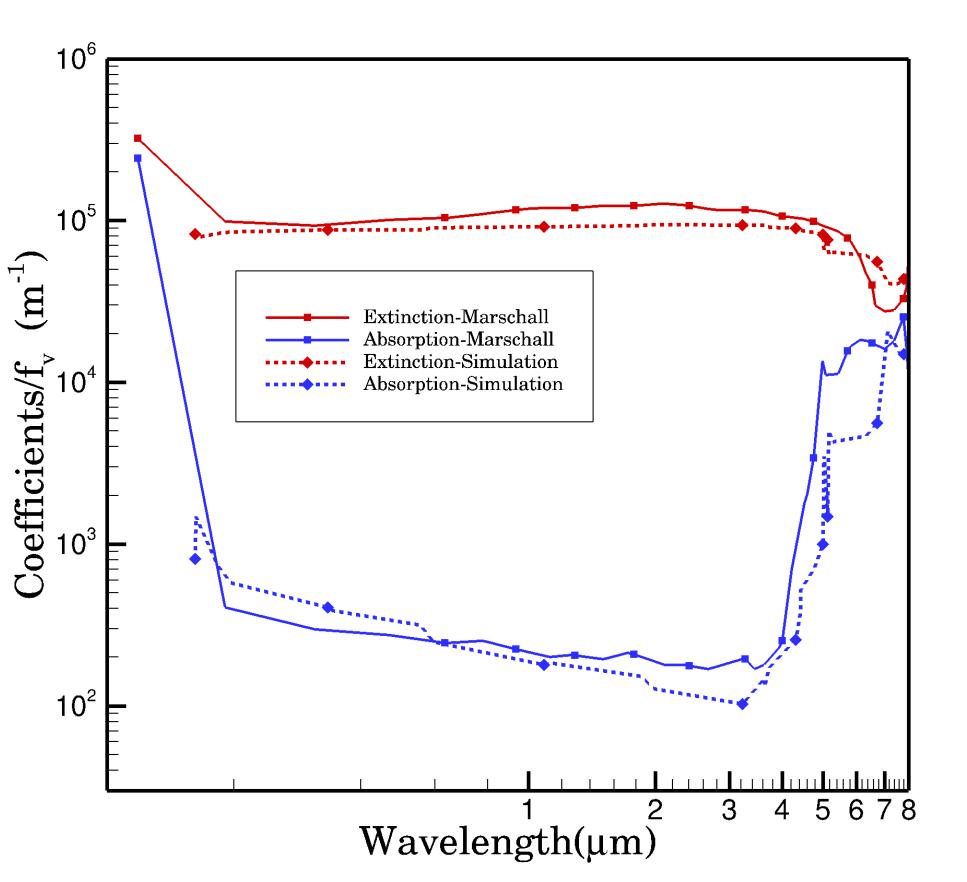
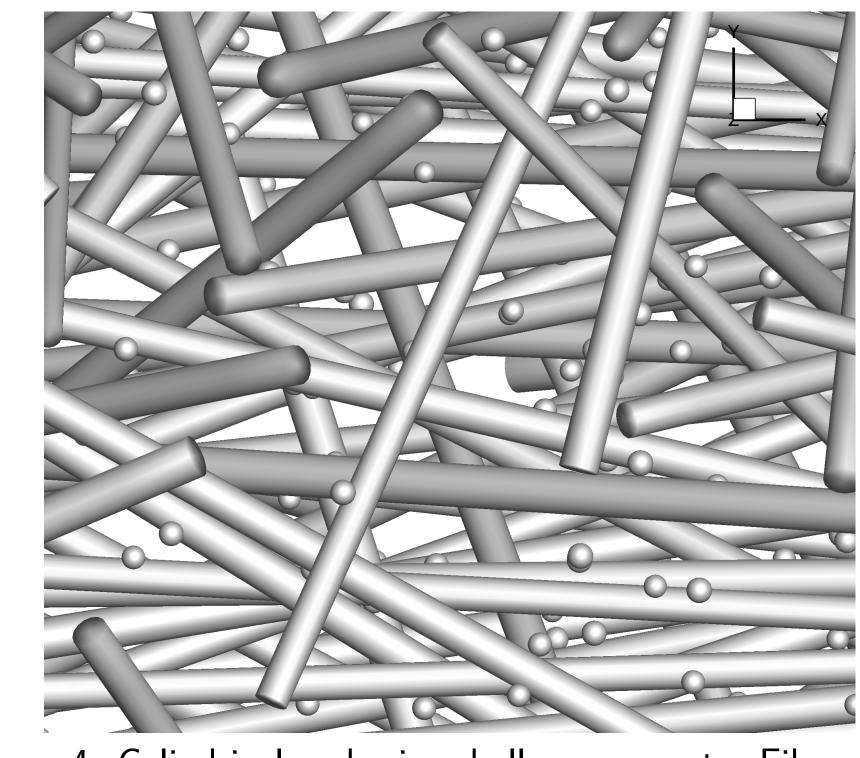


Figure: 3. Comparison of optical coefficients with Marschall

With the validation of the Sparta-rMC code, we will generate optical properties for various heat shield materials. Fig.4 shows the microstructure of cylinders and micro-balloons. These optical properties will serve as an input to the solve the Radiative Transport Equation from which the penetration depth will be determined.



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[5]	S. ter	



Conclusion

Figure: 4. Cylindrical and micro-balloon geometry Fibergen [3]

References

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